

**What's Between
Stars?**

The background features six circles of varying shades of light purple. Four are solid, and two are hollow. They are arranged in a scattered pattern around the central text. One solid circle is at the top left, another at the top right, a third at the bottom left, and a fourth at the bottom center. The two hollow circles are positioned at the top center and bottom right.

Interstellar Gas

- *There is gas between stars of night.*
- *Some of it dark, some of it bright...*



Phase of ISM

- Interstellar gas – often called ***Interstellar Medium*** (ISM) exists in different phases.
 - ***Coronal gas***: very rarefied, $T > 1$ million K; a tiny fraction by mass, about 20-50% by volume.
 - ***Diffuse interstellar gas***: several atoms per cubic cm, $T \sim 10,000$ K; about 50% by mass, 50-80% by volume.
 - ***Molecular gas***: made out of H_2 molecules; dense ($>1,000$ molecules per cm^3), $T \sim 10-30$ K; about 50% by mass, a tiny fraction by volume.

Coronal Gas

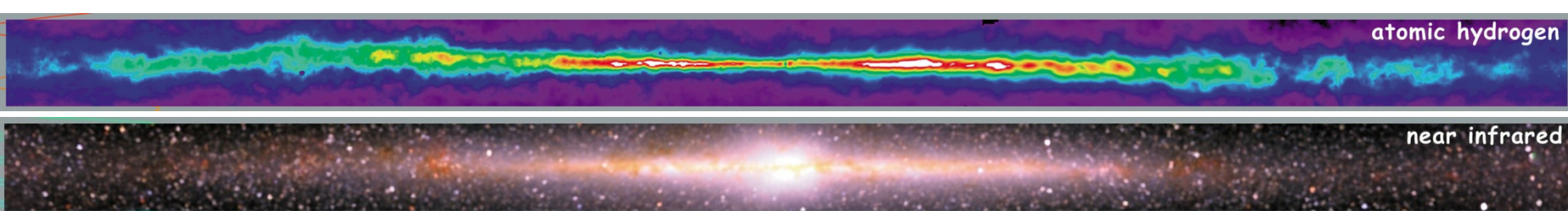
- Very tenuous, so very hard to see. Result of supernova explosions.
- We actually do not know very well how much volume it takes.
- Essentially irrelevant.



Model, not a real photograph!

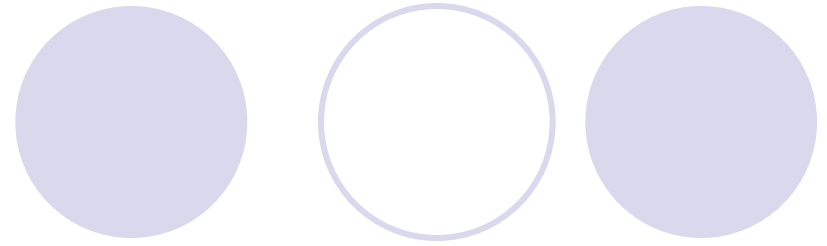
Diffuse Interstellar Gas (DIG)

- The main gas component in the Galaxy.
- Detected by the 21cm radio line of hydrogen. This is a great line to work with, since the whole universe is transparent to it!

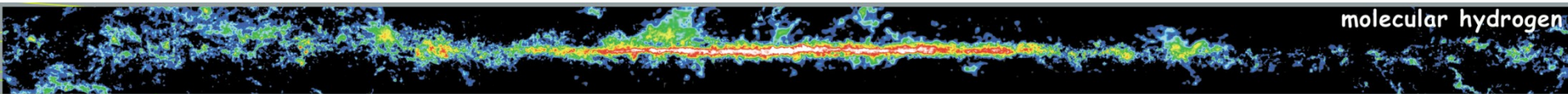


- The diffuse gas is mostly located in the disk, close to the central plane (***the plane of the Galaxy***). The gaseous disk is even narrower than the stellar disk (except on the outskirts).

Molecular Gas



- More than 99% of it is H_2 (molecular hydrogen).
- Most of the molecular gas is very dense – it has densities of 100 – 10,000 molecules per cm^3 . This is about 10 – 100 lower than the density of the best laboratory vacuum.



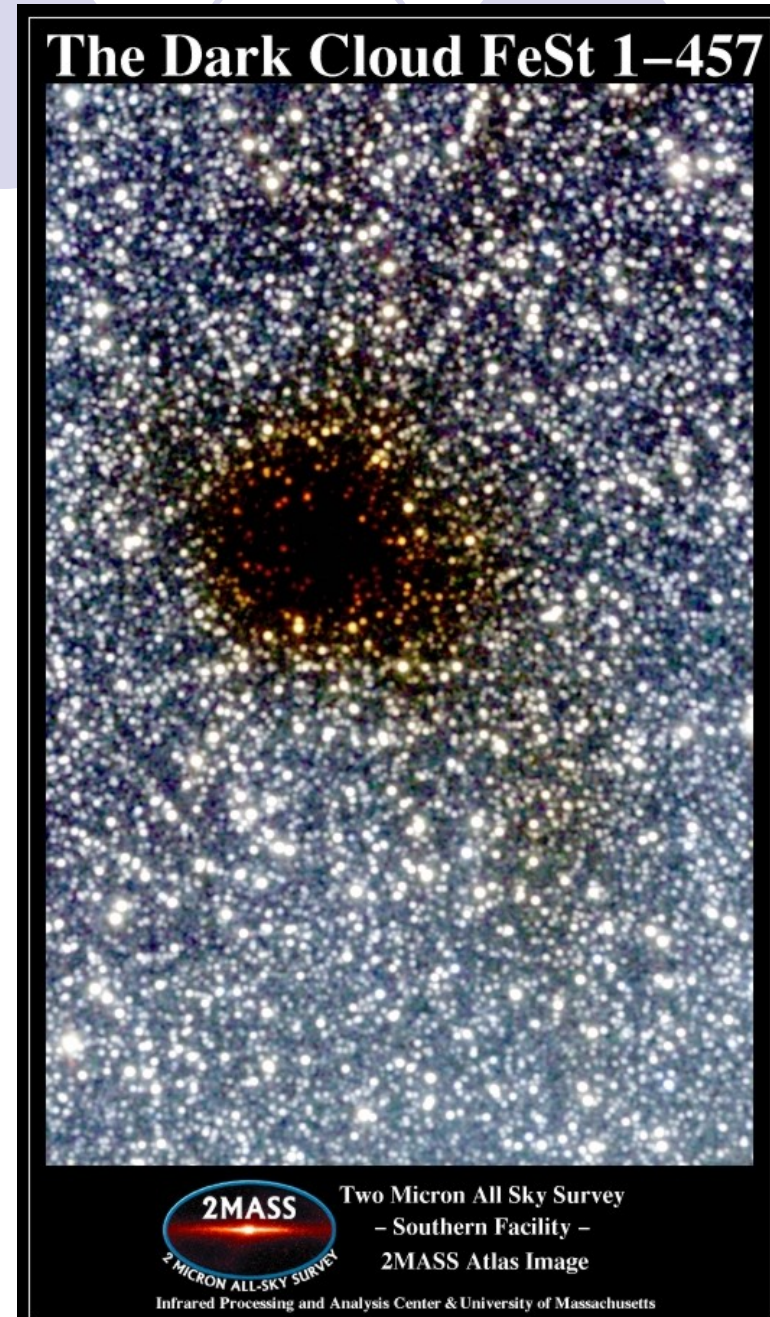
- The molecular gas is very close to the Galactic plane – the thickness of the molecular disk is only 70 pc (atomic gas ~ 150 pc, stars ~ 300 pc).

Molecular Hydrogen

- Fortunately for astronomers, hydrogen molecules are very simple, and only absorb UV radiation. They are absolutely transparent in the visible light.
- If they were not, we would not be able to see very far along the galactic plane in the visible light.
- Oops! Have I gotten it wrong?

Dark Clouds

- Molecular gas comes in separate ***dark clouds***. These clouds are dark, because they absorb most of the visible light.
- First detected by Johannes Hartmann (1865 - 1936) in 1904 using spectroscopy.
- He detected it using...
 - **A**: emission lines.
 - **B**: absorption lines.



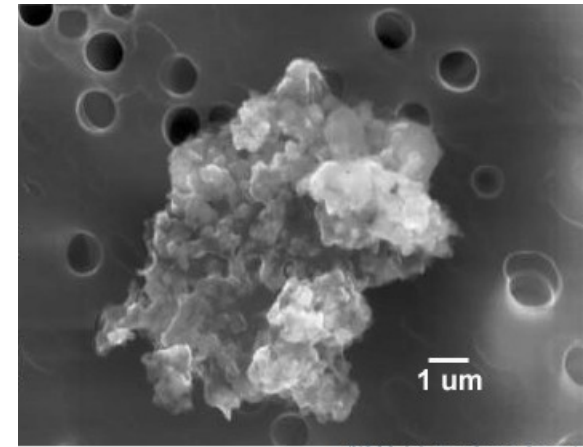
Why Are H₂ Clouds Dark?



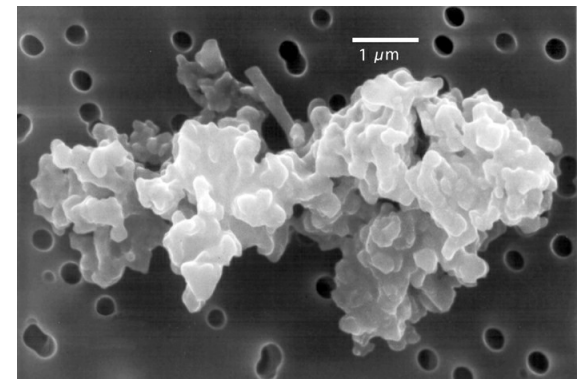
- **A:** Because molecular hydrogen absorbs the visible light very efficiently.
- **B:** Because they are located at night.
- **C:** Because there are no stars inside to light them up.
- **D:** Because molecular clouds are very cold.
- **E:** Because they contain something else that absorbs the visible light.

Cosmic Dust: Astronomer's Bane

- All molecular clouds contain cosmic dust. Dust is made of particles of various sizes, from a few molecules to sand-like grains.
- Dust particles have random shapes.
- Most of the dust is very cold: 10 – 50K (-440 – -370°F).

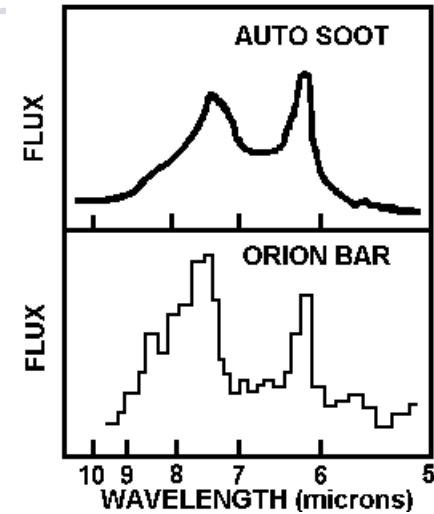


(NASA Johnson Space Center)



Dust Properties

- Chemically, these particles belong to 2 distinct types:
 - Silicates**: sand ($\text{Si}, \text{O}, \text{H}$)
 - Graphites**: soot ($\text{C}, \text{O}, \text{H}$)
- Light absorption:



Complete

None

Wavelength
(meters)

Radio

Microwave

Infrared

Visible

Ultraviolet

X-ray

Gamma Ray

10^3

10^{-2}

10^{-5}

$.5 \times 10^{-6}$

10^{-8}

10^{-10}

10^{-12}

Abundance of Cosmic Dust

- In the Milky Way, dust makes about 1% of the mass of molecular clouds. What about other galaxies?

	<i>LMC</i>	<i>SMC</i>
● Size:	8 kpc	3 kpc
● Total mass:	$2.0 \times 10^{10} M_{\odot}$	$6.0 \times 10^9 M_{\odot}$
● Disk mass:	$4 \times 10^9 M_{\odot}$	$1 \times 10^9 M_{\odot}$
● Rotation speed:	70 km/s	50 km/s
● Metallicity:	0.5	0.2

- What fraction of molecular clouds is in dust in LMC?
 - ☐ **A:** 2%
 - ☐ **B:** 1%
 - ☐ **C:** 0.5%
 - ☐ **D:** none

Why Is Plume Red?



- **A:** Because redwood is being burned.
- **B:** Because fire is red, it paints the smoke red too.
- **C:** Because soot absorbs blue light more than red.
- **D:** Because it is a sunset, and the Sun is red at sunset.
- **E:** Because the smoke is hot, it glows red.

Formation of Cosmic Dust

- Dust forms by molecules and smaller dust particles agglomerating together in collisions.
- To agglomerate efficiently, the gas has to have temperature of 1,000 – 3,000 K.
- At $T = 5,000$ K the collisions are so hard, the dust particles are destroyed instead.
- Where does cosmic dust form?
 - **A:** Molecular clouds
 - **B:** Supernova explosions
 - **C:** Atmospheres of Red Giants
 - **D:** Diffuse interstellar gas

Galactic Ecology



- Why are there all these phases? There must be a reason that gas takes all these different forms.
- What matters for gas? Unless it rotates very fast (and none of the interstellar gas does), only gravity and pressure matter.
- ***Gravity + Pressure = Density + Temperature***
- Temperature is determined by ***heating*** and ***cooling*** of the gas.
- If gas cools, it gets denser, if it heats up, it gets rarefied.

How Does Gas Cool?

- As gas cools, it gets colder, so it loses energy. But shouldn't the energy be conserved?
 - **A:** No, the conservation of energy does not apply to gas.
 - **B:** The energy is radiated away.
 - **C:** Thermal energy goes into kinetic: colder gas moves faster (winds are stronger in winter).
 - **D:** Thermal energy is converted into gravitational energy as the gas cloud collapses.

Ups and Downs in Cosmic Gas Life

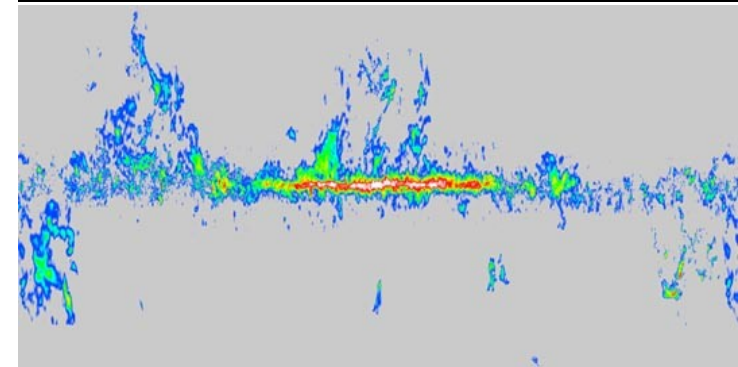
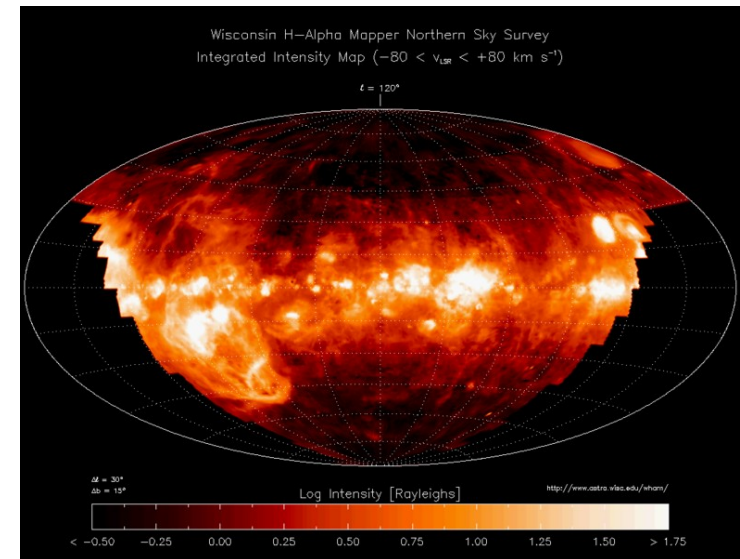
- Gas cools by glowing – i.e. emitting radiation.
- **Coronal gas:** emits X-rays.



- **Diffuse gas:** emits visible light and near IR.



- **Molecular gas:** emits far IR and microwave.



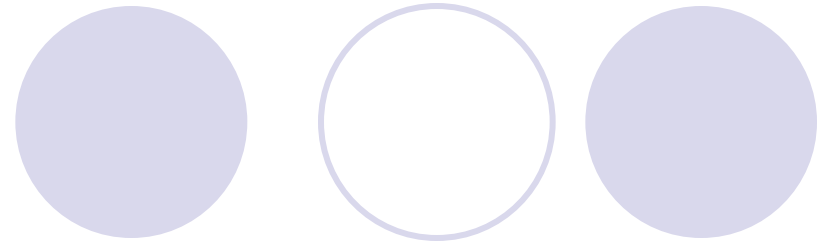
Where Does Molecular Gas Go?

- As the gas in molecular clouds keeps cooling, it gets denser, and denser, and denser...
- What happens to it eventually?
 - **A:** It becomes liquid and then solid.
 - **B:** It keeps getting denser all the time, never stops.
 - **C:** Eventually, it forms a heating source inside and stops cooling.
 - **D:** Eventually, when it gets really dark and cannot radiate anything, it stops cooling.



The danger of winter sports, or what happens when molecular gas is too careless with cooling...

Star Formation

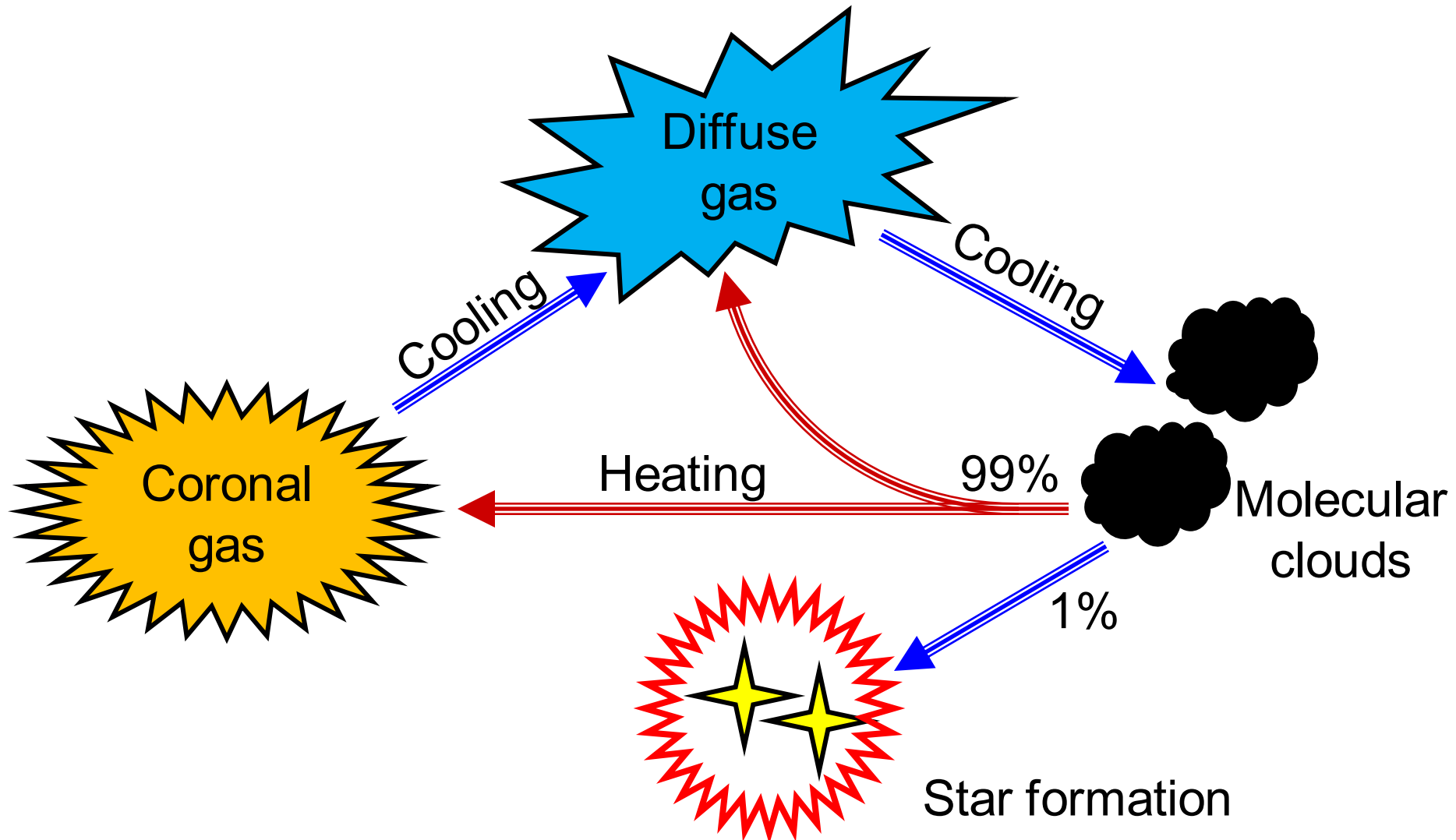


- Stars form in molecular clouds (and from molecular gas).
- “What do stars do best?” – you know that!
- Only 1% of gas needs to go into stars, for stars to be able to boil the rest of the cloud away.



- But this is a topic for another story...

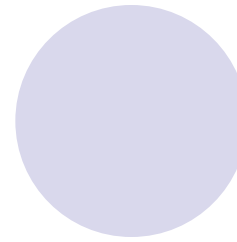
Closing the Loop



Painting by Stars

What do you see?

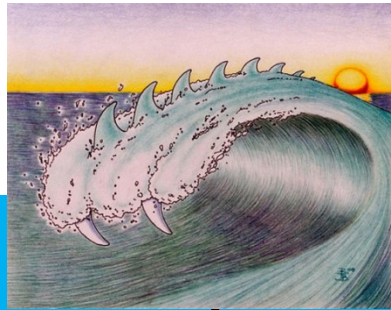
- Spiral arms.
- Cold dust is located along the inner edge.
- Heated dust is just outside it.
- Young (=blue) stars are outside hot dust.
- Diffuse gas is even further.



Spiral Density Waves

- Spiral arms are ***density waves***: they are not static objects, gas flows through them in a cycle of galactic ecology:

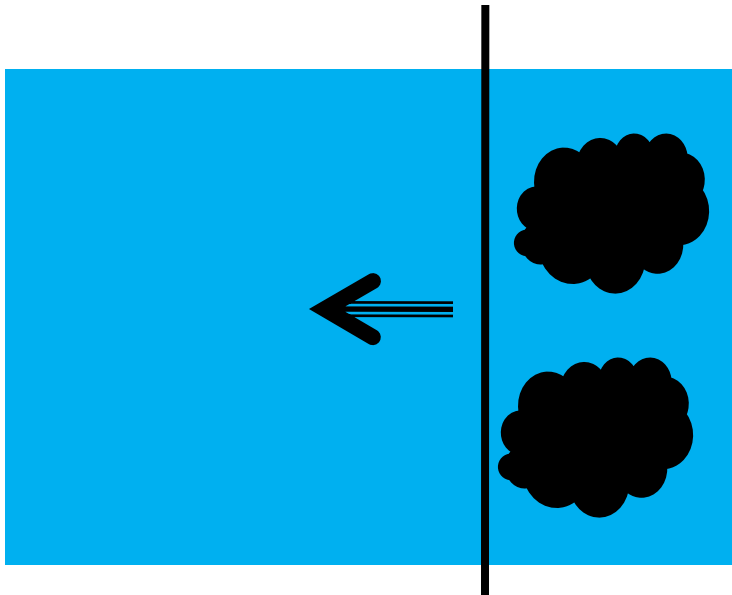
A. *Diffuse gas* gets hit by a wave...



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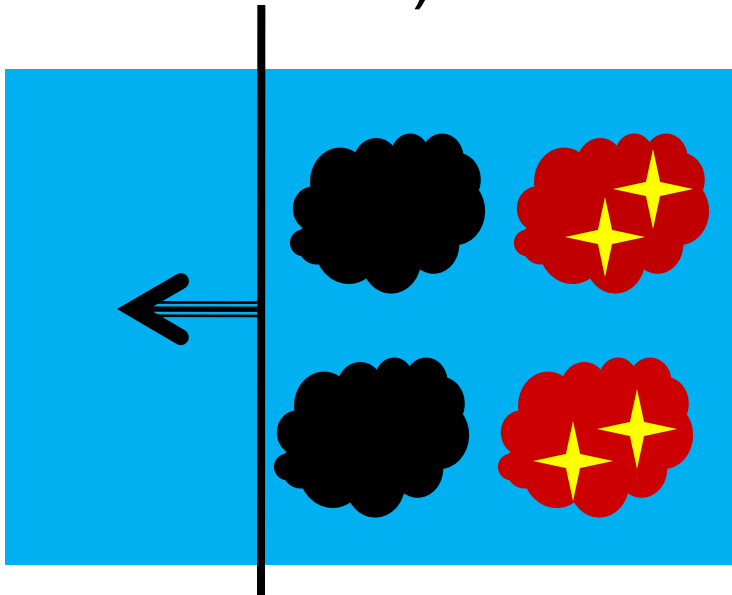
B. ...compresses, cools down, and forms cold ***molecular clouds***;



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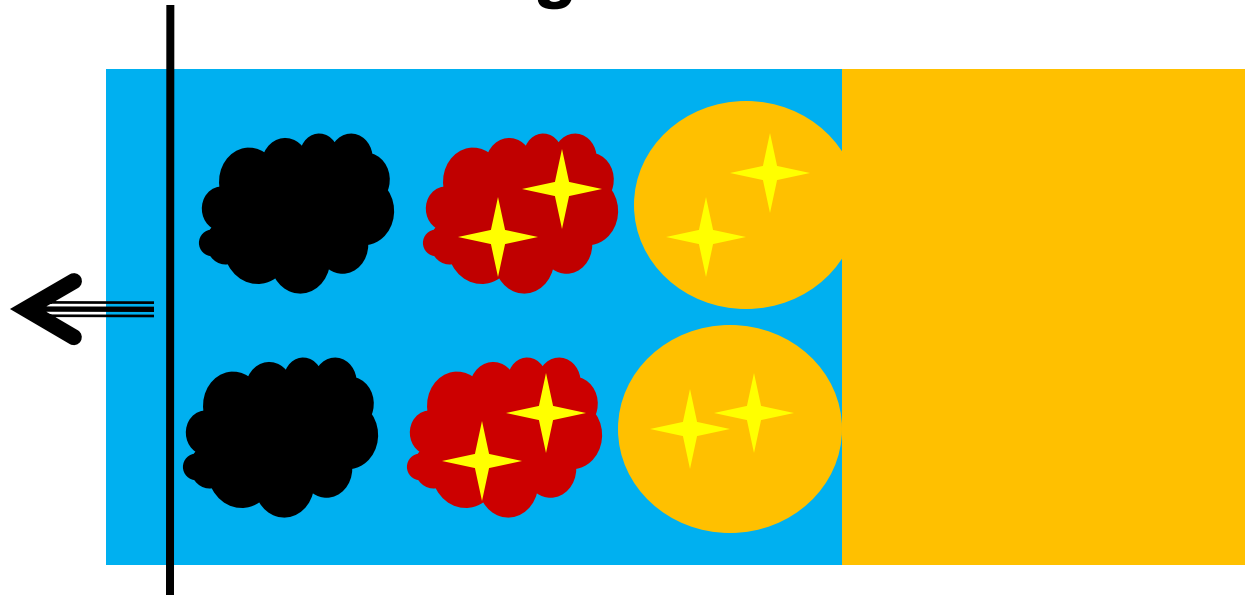
C. ***Stars*** begin to form in molecular clouds, heating the gas and dust (but the wave goes on and on...)



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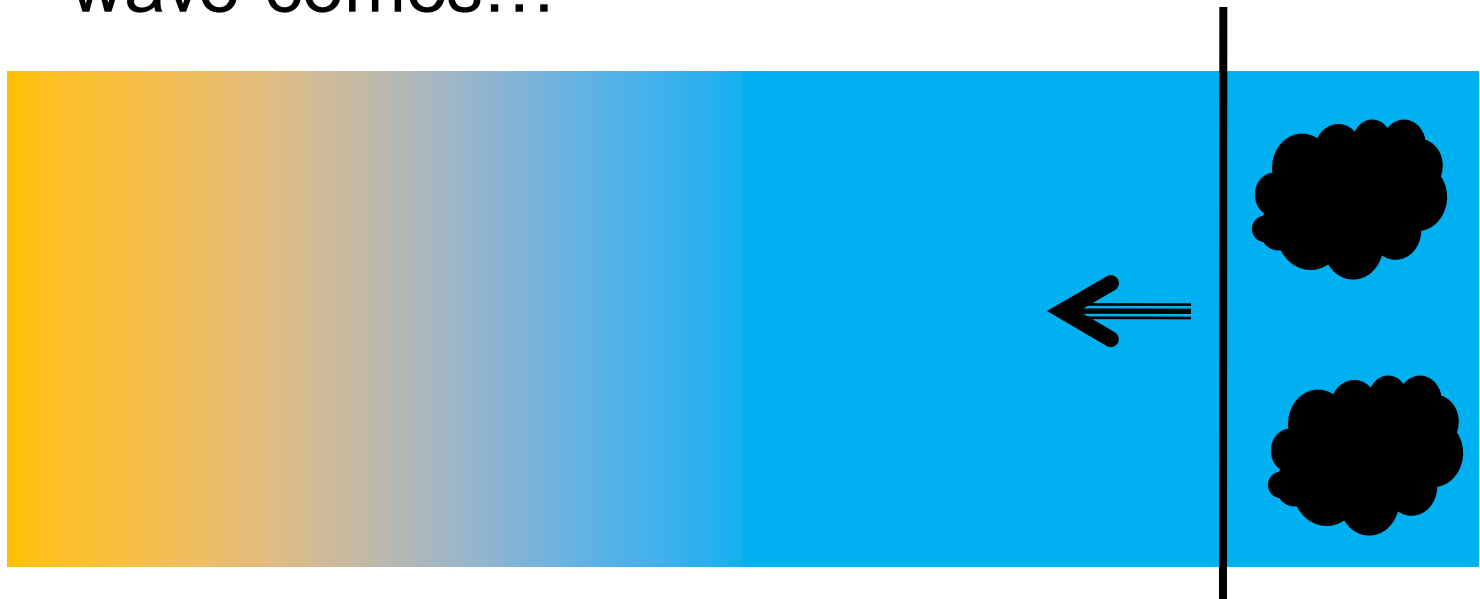
D. Eventually, molecular gas gets heated by UV radiation and supernova explosions, turning into ***coronal gas***.



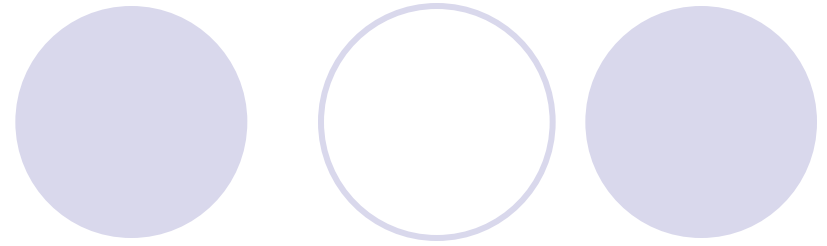
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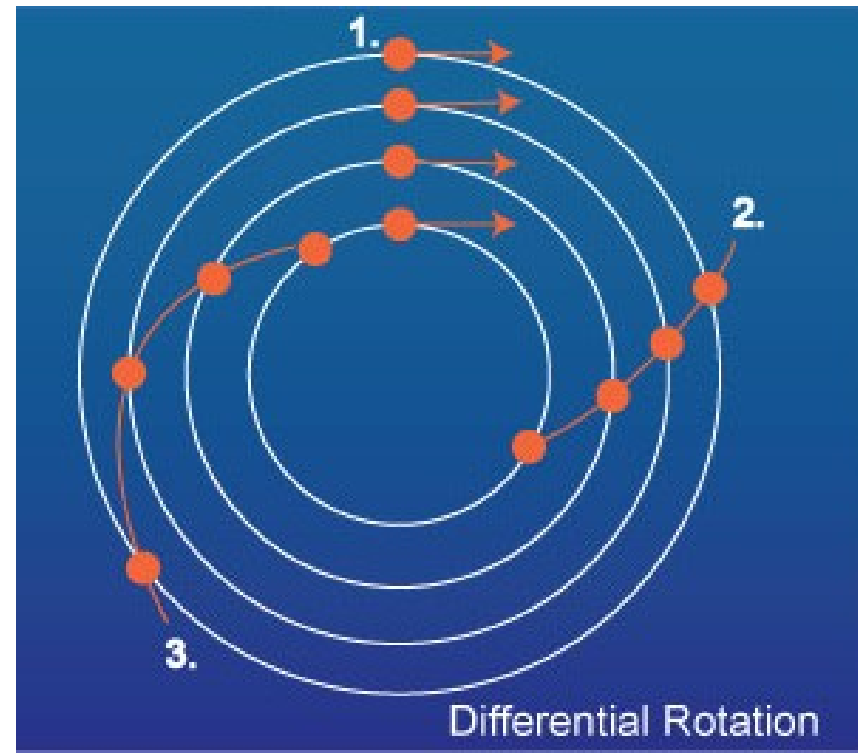
E. After the wave passed, coronal gas gradually cools into ***diffuse gas***, until the next spiral wave comes...



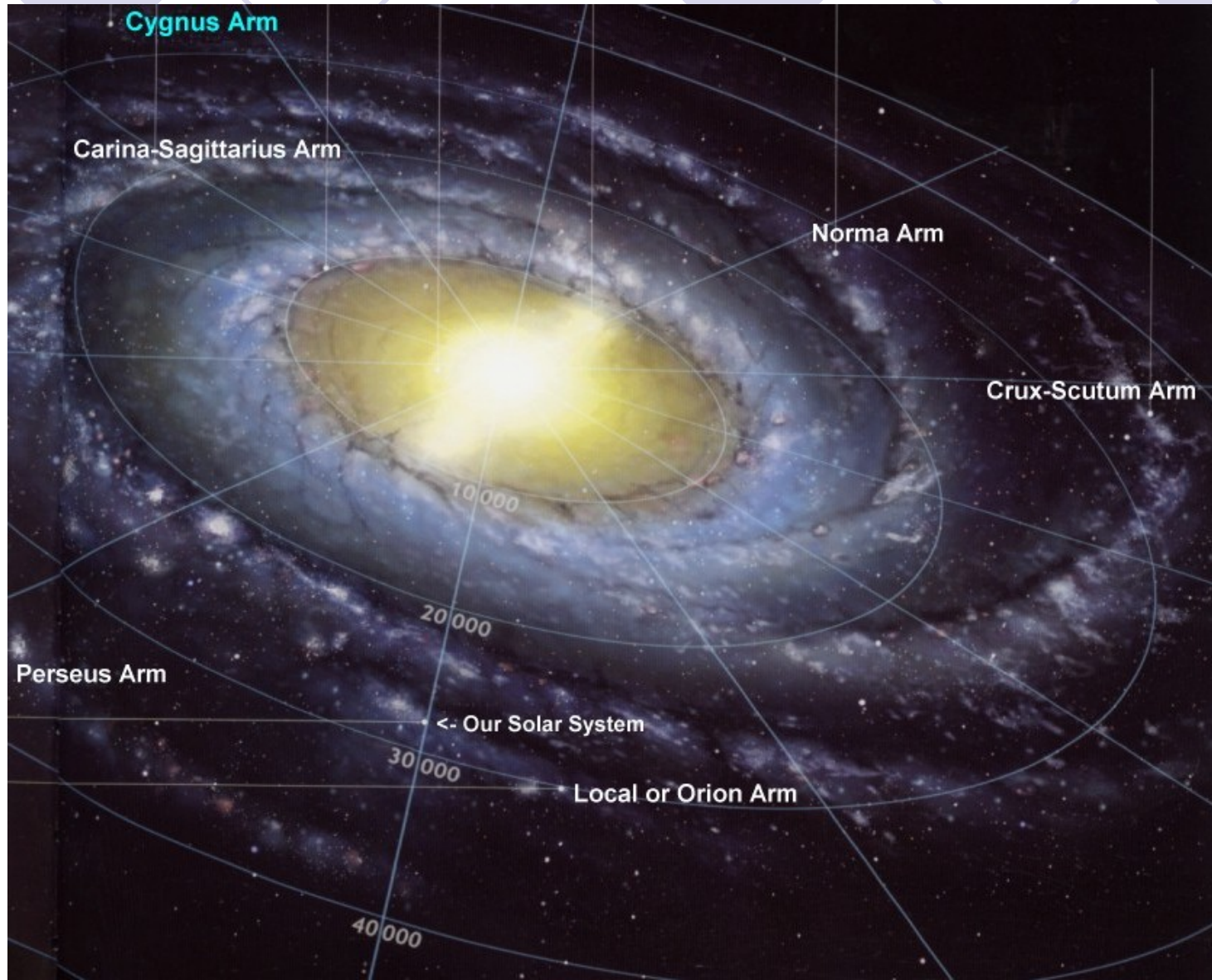
Why a Spiral?



- Recall, that ***differential rotation*** tends to stretch any pattern into a spiral.
- Rotation curve is flat, so stars closer to the center take less time to go around a smaller circle.
- A density wave gets ahead in the central region of a galaxy and falls behind on the outside.

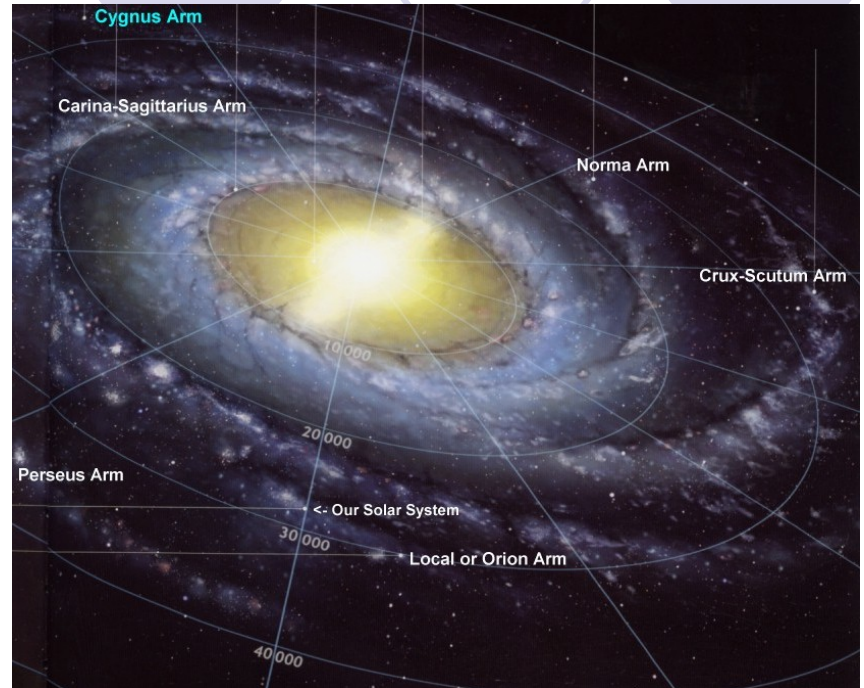


Our Spiral Arms



Which spiral arm did the Sun formed in?

- **A** Orion
- **B** Perseus
- **C** Normal
- **D** Carina-Sagittarius
- **E** Crux-Scutum
- **F** Cygnus



The Grandest Idea in Astronomy

- Waves in the ocean are crested by foam.
- Waves in a galaxy are crested by young stars!

